

Plasma as fluid

- starting from Boltzmann (Vlasov) equation, introducing "ad-hoc" force terms
- separately for electrons and ions
- $\alpha = \{e, i\}$... species

Double-fluid model

- electron (e) and ion (i) fluids "separated"

$$\frac{\partial n_\alpha}{\partial t} + \nabla \cdot (n_\alpha \vec{u}_\alpha) = 0$$

$$m_\alpha n_\alpha \left[\frac{\partial \vec{u}_\alpha}{\partial t} + (\vec{u}_\alpha \cdot \nabla) \vec{u}_\alpha \right] = q_\alpha n_\alpha (\vec{E} + \vec{u}_\alpha \times \vec{B}) - \nabla p_\alpha + \text{other forces (gravity, ...)}$$

$$p_\alpha = C_\alpha (m_\alpha n_\alpha)^{\gamma_\alpha} \dots \text{E.O.S.}$$

γ_α ... ratio of specific heats

+ Maxwell

$$\epsilon_0 \nabla \cdot \vec{E} = \sum_\alpha n_\alpha q_\alpha$$
$$\nabla \times \vec{E} = - \frac{\partial \vec{B}}{\partial t}$$

coupling here, so fluids are not that independent

$$\nabla \cdot \vec{B} = 0$$

$$\frac{1}{\mu} \nabla \times \vec{B} = \sum_\alpha n_\alpha q_\alpha \vec{u}_\alpha + \epsilon_0 \frac{\partial \vec{E}}{\partial t}$$

→ sum: equations 2 · (1 + 3 + 1) + 1 + 3 + 1 + 3 = 16
variables 16

$$\text{but! } 0 = \nabla \cdot (\nabla \times \vec{E}) = - \nabla \cdot \frac{\partial \vec{B}}{\partial t} = - \frac{\partial}{\partial t} \nabla \cdot \vec{B}$$

$$\Rightarrow \nabla \cdot \vec{B} = \text{const}$$

similarly for $\nabla \times \vec{B} \Rightarrow$ scalar Maxwell are not independent ("initial conditions")

$\Rightarrow$ only 16 independent eqs.

Note: plasma approximation

On large scales quasi-neutrality

$$\Rightarrow n_e = n_i ; \text{ do not use } \nabla \cdot \vec{E} = \sum n_\alpha q_\alpha$$

Highly ionised plasma

capturing particle interactions on scales smaller than Debye: "collision" $\rightarrow$ need to parameterize

$$\text{eg. } m_e n_e \frac{d\vec{u}_e}{dt} = -en_e (\vec{E} + \vec{u}_e \times \vec{B}) - \nabla p_e + \vec{P}_{ei}$$

momentum exchange between e and i
momentum conserved = symmetry
 $\vec{P}_{ei} = -\vec{P}_{ie}$

heuristics

efficiency of collisions
 $\Rightarrow$ specific resistivity

$$P_{ei} = \underbrace{\gamma}_{\text{efficiency of collisions}} \underbrace{e^2}_{\text{Coulomb interaction}} \underbrace{n_e}_{\Rightarrow \text{charge}^2} \underbrace{n_i (\vec{u}_i - \vec{u}_e)}_{\text{relative speeds}}$$

how often do e^- meet i^+

proportional to electron flux

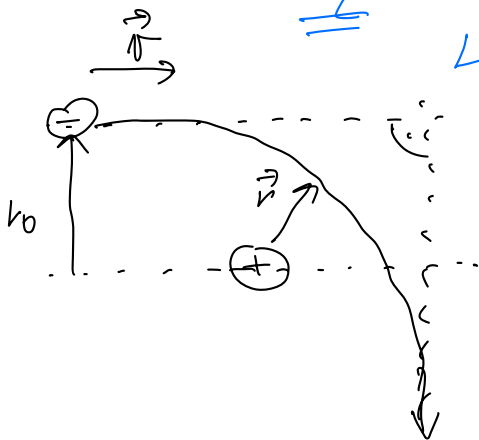
alternatively

$$P_{ei} = \underbrace{m_e n (\vec{u}_i - \vec{u}_e)}_{\text{momentum}} \underbrace{v_{ei}}_{\text{momentum per time of collisions}} \quad \text{collisional frequency} \sim \frac{1}{\tau_{ei}}$$

compare: $v_{ei} = \frac{ne^2}{m_e} \gamma = \frac{\omega_p^2 \epsilon_0 \gamma}{\omega_p^2} = \frac{ne^2}{m_e \epsilon_0}$ plasma frequency

γ is a key

$\rightarrow$ simplistic model



$$\text{Coulombic: } F = -\frac{e^2}{4\pi\epsilon_0 r^2}$$

interaction at the closest point takes $\Delta t \sim \frac{r_0}{v}$... estimate

"interaction is important only at close distances"

$$\Rightarrow \text{momentum exchange: } \Delta(m_e v) \sim |F \Delta t| = \frac{e^2}{4\pi\epsilon_0 r_0 v}$$

assumption: deflection by $90^\circ \Rightarrow$ all momentum lost

$$\Rightarrow \Delta(u_e v) = m_e v = \frac{e^2}{4\pi\epsilon_0 v} \quad \text{looking for:}$$

$$r_{90} = \frac{e^2}{4\pi\epsilon_0 m_e v^2}$$

effective cross-section: $\sigma = \pi r_{90}^2 = \frac{e^4}{16\pi\epsilon_0^2 m_e^2 v^4}$

an collisional frequency: $\nu_{ei} = \sigma n v = \frac{n e^4}{16\pi\epsilon_0^2 m_e^2 v^3}$

now relate electron velocity to temperature, assuming the equilibrium

$m_e v_e^2 \sim k_B T_e$ and previously: $\nu_{ei} = \frac{n e^4}{m_e} \zeta$

hence $\zeta = \frac{\nu_{ei} m_e}{n e^2} = \frac{e^2}{16\pi\epsilon_0^2 m_e v^3} = \frac{e^2 m_e^{1/2}}{16\pi\epsilon_0^2 (k_B T_e)^{3/2}}$

So far - only for a specific $v_0 = v_{90}$

We should account for all, integrating over $\int_0^\infty \frac{1}{r} dr$, but it diverges:

So limits: lower $\rightarrow$ perpendicular deflection
upper $\rightarrow$ Debye length

$$\Rightarrow \int_{r_{90}}^{\lambda_D} \frac{1}{r} dr = \left[\ln r \right]_{r_{90}}^{\lambda_D} = \ln \frac{\lambda_D}{r_{90}} \equiv \underbrace{\ln \Lambda}_{\text{Coulombic logarithm}}$$

Hence $\zeta = \frac{e^2 m_e^{1/2}}{16\pi\epsilon_0^2 (k_B T_e)^{3/2}} \ln \Lambda$

! $\zeta \neq \zeta(u_e)$

$\zeta \propto T^{-3/2} \Rightarrow T \uparrow, \zeta \rightarrow$ Joule heating not efficient in plasma

for $k_B T_e \sim 100 \text{ eV} \dots \zeta = 5 \times 10^{-7} \Omega m \rightarrow$ like copper

Note: $P_{ei} = \zeta e^2 n^2 (\vec{u}_i - \vec{u}_e) = e n_e E$; $E = \zeta e n (u_i - u_e) = \zeta \vec{j} \rightarrow$ Ohm's law

Single fluid model

→ we cannot distinguish between electrons and ions, so a mixture is considered

→ combined variables are closer to be measured

Definition:

plasma approximation

bulk density: $\rho \equiv n_i M + n_e m \sim n(m+M) \sim nM$
 ↑ usually $m \ll M$

bulk velocity: $\vec{w} \equiv \frac{1}{\rho} (n_i M \vec{u}_i + n_e m \vec{u}_e)$
 $\sim \frac{n(M \vec{u}_i + m \vec{u}_e)}{n(M+m)} = \frac{M \vec{u}_i + m \vec{u}_e}{M+m}$

↑ plasma approx.

single-charged ions

bulk current density: $\vec{j} \equiv e(n_i \vec{u}_i - n_e \vec{u}_e) \sim ne(\vec{u}_i - \vec{u}_e)$

bulk pressure: $p = p_i + p_e$

bulk charge density: $\rho_e = n_i q_i + n_e q_e = e(n_i - n_e)$

! new variables are linear combinations of the old ones:

$(n_e, n_i) \rightarrow (\rho, \rho_e); (\vec{u}_e, \vec{u}_i) \rightarrow (\vec{w}, \vec{j})$

one exception: pressure

Linear combinations of fluid equations

$\left. \begin{aligned} \frac{\partial n_i}{\partial t} + \nabla \cdot (n_i \vec{u}_i) &= 0 & \cdot M \\ \frac{\partial n_e}{\partial t} + \nabla \cdot (n_e \vec{u}_e) &= 0 & \cdot m \end{aligned} \right\} +$ $\frac{\partial}{\partial t} (M n_i + m n_e) + \nabla \cdot [M n_i \vec{u}_i + m n_e \vec{u}_e] = 0$ $\frac{\partial p}{\partial t} + \nabla \cdot (p \vec{w}) = 0$	$\left. \begin{aligned} \frac{\partial n_i}{\partial t} + \nabla \cdot (n_i \vec{u}_i) &= 0 & \cdot e \\ \frac{\partial n_e}{\partial t} + \nabla \cdot (n_e \vec{u}_e) &= 0 & \cdot e \end{aligned} \right\} -$ $\frac{\partial}{\partial t} (e n_i - e n_e) + \nabla \cdot (e n_i \vec{u}_i - e n_e \vec{u}_e) = 0$ $\frac{\partial \rho_e}{\partial t} + \nabla \cdot \vec{j} = 0$
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$$\begin{aligned}
 Mn \frac{\partial \vec{w}_i}{\partial t} &= en(E + \vec{w}_i \times \vec{B}) - \nabla p_i + Mng + P_{ei} \\
 mn \frac{\partial \vec{w}_e}{\partial t} &= -en(E + \vec{w}_e \times \vec{B}) - \nabla p_e + mng + P_{ei}
 \end{aligned}
 \quad \left. \vphantom{\begin{aligned} Mn \frac{\partial \vec{w}_i}{\partial t} \\ mn \frac{\partial \vec{w}_e}{\partial t} \end{aligned}} \right\} +$$

$\uparrow$ convective derivative neglected $\uparrow$ "ad-hoc" gravity collisions, but $P_{ei} = -P_{ei}$

$$\underbrace{n(M+m)}_J \frac{\partial}{\partial t} \left(\underbrace{\frac{M\vec{w}_i + m\vec{w}_e}{M+m}}_{\vec{w}} \right) = \underbrace{en(\vec{w}_i - \vec{w}_e)}_{\vec{j}} \times \vec{B} - \nabla p + \underbrace{n(M+m)}_J \vec{g}$$

$$\underbrace{J}_{\downarrow} \frac{\partial \vec{w}}{\partial t} = \vec{j} \times \vec{B} - \nabla p + J\vec{g}$$

(generalised to $J(\frac{\partial \vec{w}}{\partial t} + (\vec{w} \cdot \nabla)\vec{w}) = \dots$)

note: $\vec{j} \parallel \vec{w} \rightarrow$ do not have to be parallel

different combination $\rightarrow$ multiple by mass of the other species:

$$\begin{aligned}
 mMn \frac{\partial \vec{w}_i}{\partial t} &= eMn(E + \vec{w}_i \times \vec{B}) - M\nabla p_i + mMng + mP_{ei} \\
 mMn \frac{\partial \vec{w}_e}{\partial t} &= -eMn(E + \vec{w}_e \times \vec{B}) - M\nabla p_e + mMng + MP_{ei}
 \end{aligned}
 \quad \left. \vphantom{\begin{aligned} mMn \frac{\partial \vec{w}_i}{\partial t} \\ mMn \frac{\partial \vec{w}_e}{\partial t} \end{aligned}} \right\} -$$

$$\begin{aligned}
 nmM \frac{\partial}{\partial t} (\vec{w}_i - \vec{w}_e) &= en(M+m)E + en(m\vec{w}_i + M\vec{w}_e) \times \vec{B} - \\
 &\quad - M\nabla p_i + M\nabla p_e - (M+m)P_{ei} \quad \left| : en(M+m) = eJ \right.
 \end{aligned}$$

$$m\vec{w}_i + M\vec{w}_e = \underbrace{M\vec{w}_i + m\vec{w}_e}_{\text{add and remove } m\vec{w}_e \text{ and } M\vec{w}_i} + M(\vec{w}_e - \vec{w}_i) + m(\underbrace{\vec{w}_i - \vec{w}_e}_{\vec{j}/ne}) =$$

$$= \frac{J}{n} \vec{w} - (M-m) \frac{\vec{j}}{ne}$$

$$\text{also } \vec{P}_{ei} = \frac{1}{2} e^2 n^2 (\vec{w}_i - \vec{w}_e) = \frac{1}{2} en \vec{j}$$

$$\vec{E} + \vec{w} \times \vec{B} - \frac{1}{cJ} \left[\frac{Mmn}{e} \frac{\partial}{\partial t} \frac{\vec{j}}{n} + (M+m) \vec{j} \times \vec{B} + m\nabla p_i - M\nabla p_e \right]$$

$\rightarrow$ take $m \ll M$

$$\rightarrow \frac{Mmn}{c} \frac{\partial \dot{j}}{\partial t} = MnB \frac{m}{eB} \frac{\partial \dot{j}}{\partial t} = MnB \omega_c^{-1} \frac{\partial \dot{j}}{\partial t}$$

$$\text{if } \omega_c \gg \left(\frac{\partial \dot{j}}{\partial t} \right)^{-1}$$

$$\text{then } MnB \omega_c^{-1} \frac{\partial \dot{j}}{\partial t} \ll (M-m) \dot{j} \times B$$

$$\Rightarrow \underbrace{\vec{E} + \vec{\omega} \times \vec{B} - \eta \vec{j}}_{\text{Hall's current}} = \frac{1}{en} (\underbrace{\vec{j} \times \vec{B}}_{\text{often neglected}} - \nabla p_e) \quad \text{generalised Ohm's law}$$

Double-fluid model

$$\rho \frac{\partial \vec{\omega}}{\partial t} = \vec{j} \times \vec{B} - \nabla p + \rho \vec{g}$$

$$\vec{E} + \vec{\omega} \times \vec{B} = \eta \vec{j}$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \rho \vec{\omega} = 0$$

$$\frac{\partial p_e}{\partial t} + \nabla \cdot \vec{j} = 0$$

+ Maxwell's equations $\rightarrow$ should be solvable,
16 eq, 16 variables

$\rightarrow$ almost equivalent to double-fluid model, but
some information is lost in the process (pressure
neglectances)

Approximations to plasma fluid

Force-free and ideal MHD

force-free $\rightarrow$ plasma does "not feel" the Lorentz force

$$\vec{f} = \mu_0 \vec{E} + \vec{j} \times \vec{B} = 0$$

$\uparrow$ "effective" or "bulk" acting on mixture

$\rightarrow$ or Lorentz force small and negligible
compared to other forces

$\rightarrow$ or EM field induces forces that cancel external
forces

for E small: $\vec{j} \times \vec{B} = 0 \dots$ often used in astrophysics

for $\nabla \times \vec{B} = \mu_0 \vec{j}$

$$(\nabla \times \vec{B}) \times \vec{B} = 0 \Rightarrow \nabla \times \vec{B} \parallel \vec{B}$$

$$\Rightarrow \nabla \times \vec{B} = \alpha \vec{B} \dots \text{linear force-free}$$

may change in space

ideal MHD - in a co-moving frame E disappears

$$\vec{E}' = \vec{E} + \vec{w} \times \vec{B} = 0$$

using gen. Ohm's law: $\vec{E}' = \eta \vec{j}$, $\vec{E}' = 0$ for $\vec{j} = 0$ or $\eta = 0$

if $E \cdot B = 0$ - fields degenerated $\rightarrow$ force-free and ideal MHD are the same
- they coincide if $\vec{j} = \eta \nabla \times \vec{B}$

Frozen field

consequence of the ideal MHD

highly conductive plasmas, strong plasmas $\leftrightarrow$ field coupling

Start from Ampère's law, taking a curl

$$\nabla \times (\nabla \times \vec{B}) = \nabla \nabla \cdot \vec{B} - \Delta \vec{B} = \nabla \times \left(\mu_0 \vec{j} + \mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t} \right)$$

$\nabla \cdot \vec{B} = 0$ Ohm's law and $\sigma \equiv \frac{1}{\eta}$ conductivity

$$\epsilon_0 \mu_0 \nabla \times (\vec{E} + \vec{w} \times \vec{B}) = -\epsilon_0 \mu_0 \nabla \times \frac{\partial \vec{E}}{\partial t} - \Delta \vec{B}$$

if this is finite, for $\eta \rightarrow 0$ $\sigma \rightarrow \infty$

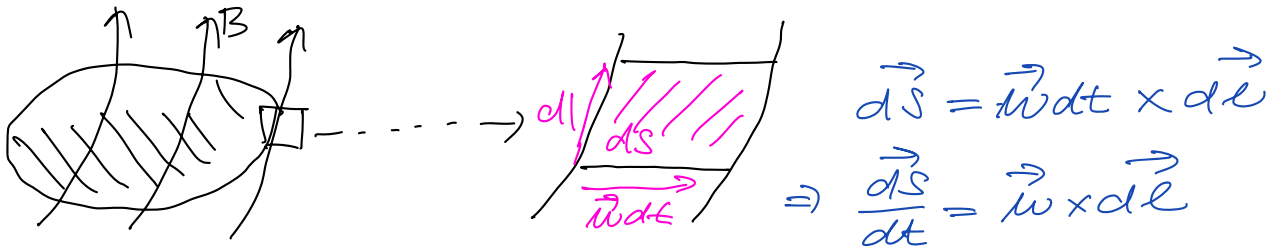
$$\Rightarrow \nabla \times (\vec{E} + \vec{w} \times \vec{B}) \rightarrow 0$$

Faraday's law

$$-\frac{\partial \vec{B}}{\partial t} + \nabla \times (\vec{w} \times \vec{B}) = 0 \Rightarrow \nabla \times (\vec{w} \times \vec{B}) = \frac{\partial \vec{B}}{\partial t} \leftarrow \text{frozen field}$$

! curl being zero does not mean $\vec{E} + \vec{w} \times \vec{B} = 0$!

interpretation: magnetic flux through closed surface does not change with time



$$d\Phi = B \cdot dS$$

$$\begin{aligned} \frac{d\Phi}{dt} &= \frac{d}{dt} \int B \cdot dS = \int B \cdot \left(\frac{dS}{dt} \right) + \int \frac{\partial B}{\partial t} \cdot dS = \text{Stokes theorem} \\ &= \oint B \cdot (u \times dl) + \int \frac{\partial B}{\partial t} \cdot dS = - \oint (u \times B) \cdot dl + \int \frac{\partial B}{\partial t} \cdot dS = \\ &= - \int \nabla \times (u \times B) \cdot dS + \int \frac{\partial B}{\partial t} \cdot dS = - \int \left[\nabla \times (u \times B) - \frac{\partial B}{\partial t} \right] \cdot dS = 0 \\ &\quad \text{Q.E.D.} \end{aligned}$$

Drifts in plasma fluid

Drifts perpendicular to B

$$\underbrace{m n \left(\frac{\partial u}{\partial t} + (u \cdot \nabla) u \right)}_{\text{large scale, slow changes}} = \underbrace{q n (E + u \times B) - \nabla p}_{\text{small scales, fast changes}}$$

$\Rightarrow$ RHS neglected

$$0 = q n (E + u \times B) - \nabla p \quad / \times B$$

$$0 = q n (E \times B + (u \times B) \times B) - \nabla p \times B =$$

$$= q n (E \times B - u B^2 + B u \cdot B) - \nabla p \times B \quad \left. \begin{array}{l} \text{perp.} \\ \text{component} \end{array} \right\}$$

$$0 = q n (E \times B - \underbrace{(u_{\perp})^2}_{\text{diamagnetic drift}} B^2) - \nabla p \times B$$

$$u_{\perp} = \frac{E \times B}{B^2} - \underbrace{\frac{\nabla p \times B}{q n B^2}}_{\text{diamagnetic drift}} = v_E + v_D$$

$v_D \perp \nabla p \Rightarrow (u \cdot \nabla) u$ negligible if $E=0$

$\nabla \Phi \parallel \nabla p \vee E = -\nabla \Phi \Rightarrow (u \cdot \nabla) u$ still negligible
otherwise $(u \cdot \nabla) u$ term complicates things

Drifts parallel to B

$$m n \left(\frac{\partial u_z}{\partial t} + (u \cdot \nabla) u_z \right) = q n E_z - \frac{\partial p}{\partial z}$$

E.O.S. $\frac{\partial p}{p} = \gamma \frac{\partial n}{n}$; $\gamma = 1 \rightarrow$ isothermic
($u \cdot \nabla$) $u_z \rightarrow 0$, u_z homogeneous

$$\Rightarrow \nabla p = k_B T \nabla n$$

For electrons: $q \rightarrow -e$, $m \rightarrow 0$, $T \rightarrow T_e$, $E_z = -\frac{\partial \phi}{\partial z}$

then $e \frac{\partial \phi}{\partial z} = \frac{k_B T_e}{n} \frac{\partial n}{\partial z}$

hence $e \phi = k_B T_e \ln n + C$

$$\Rightarrow n = n_0 \exp \left[\frac{e \phi}{k_B T_e} \right]$$

Boltzmann relation hold also for fluids

$\Rightarrow$ ions creating concentrations $\rightarrow$ electrons shield
following the Boltzmann relation